

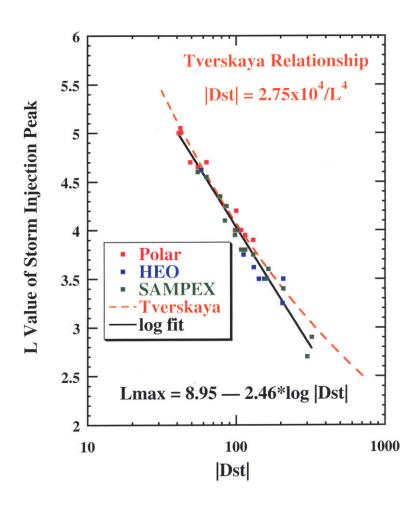
CEPPAD Science Report

- Relativistic electrons
- New opportunities for Polar CEPPAD science





Relativistic Electron Injection Location



- Excellent correlation with |Dst|
- Correlation holds for all pitch angles
- Intensity increases most rapidly at peak location (not shown)
- Intensity increases more slowly as electron energy increases (not shown)



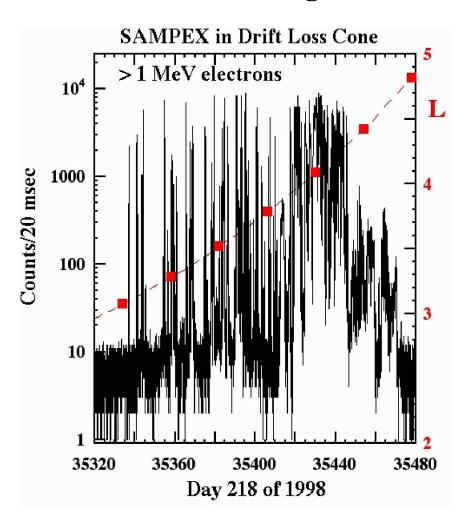
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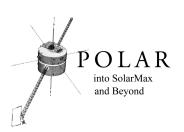
Pitch-angle Scattering Highly Bursty

3

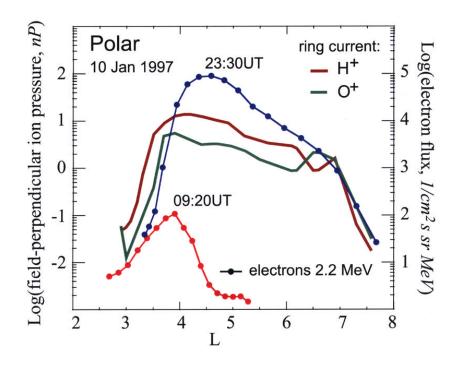


- SAMPEX sees bursts with a time structure substantially shorter than one second
- Burst seen in 25% of passes during storms
- Dawn sector favored





Example of Correlation with Polar Ring-Current Measurements



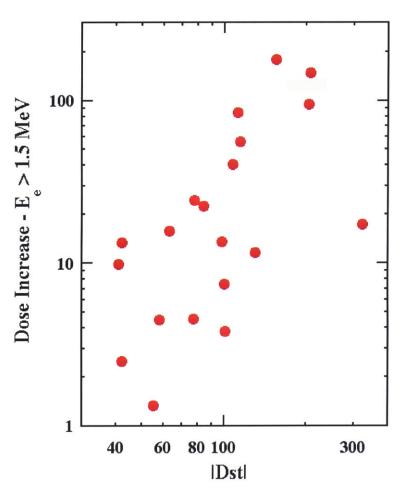
- Electron intensity peak near region of peak pressure
- Only a few such measurements analyzed to date
 - Polar infrequently in region on interest around max |Dst|
 - After electrons appear, MICS background is high



4



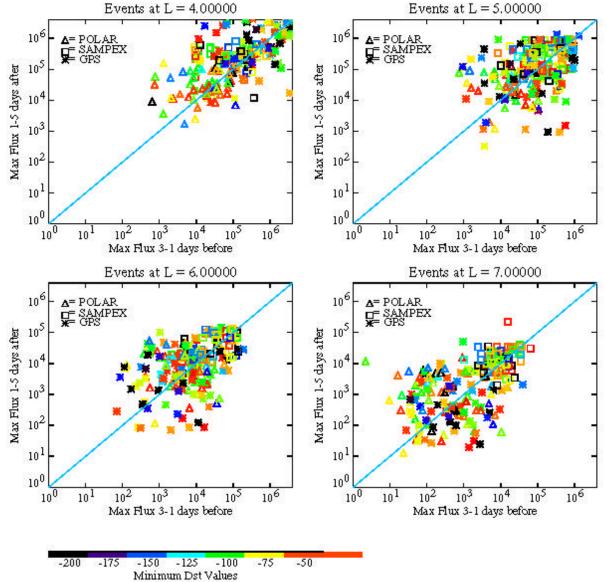
Correlation of Electron Intensity with |Dst| Poor

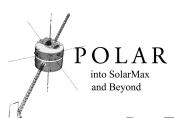


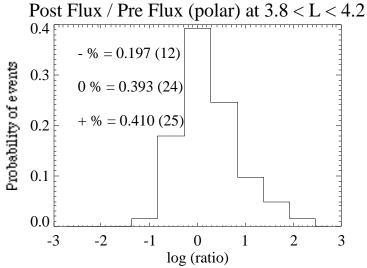
- Large scatter; largest storms in particular have not yielded largest injections
- Only ~ 50% of storms lead to relativistic electron enhancement

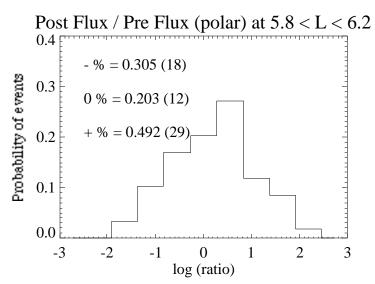


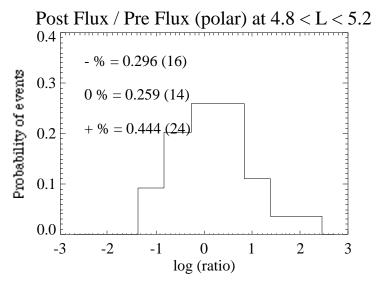


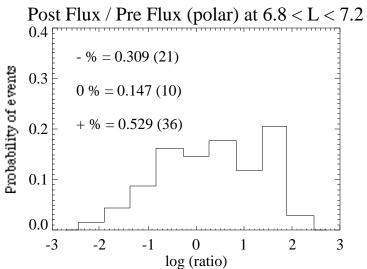














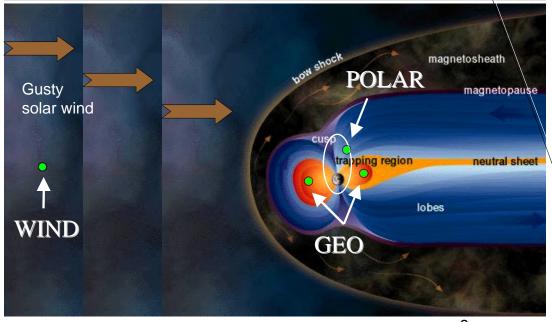
POLAR GGS Reveals Direct Connection Between Solar Wind and Magnetospheric Waves: A Global Driven Response

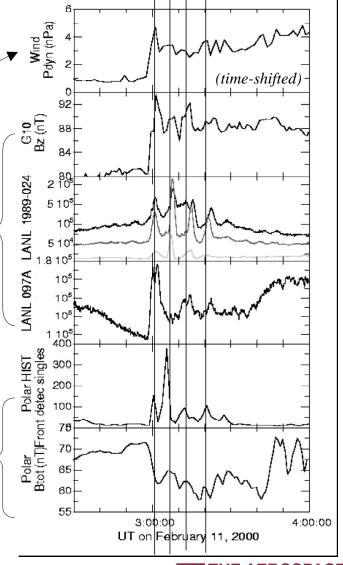
A typical event – 11 February 2000

WIND SWE observes a small shock and associated dynamic pressure waves well upstream of Earth

GOES/LANL spacecraft see ULF wave response at driver period in particles and magnetic fields

POLAR CEPPAD and MFE observations reveal coherence and extent of global oscillations





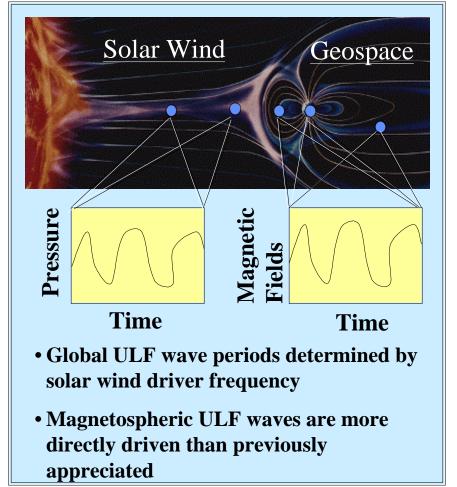


Revisiting an Old Sun-Earth Connection: Solar Wind Pressure Pulses Can Directly Drive ULF Magnetosphere Waves

Pre-ISTP Conventional Wisdom

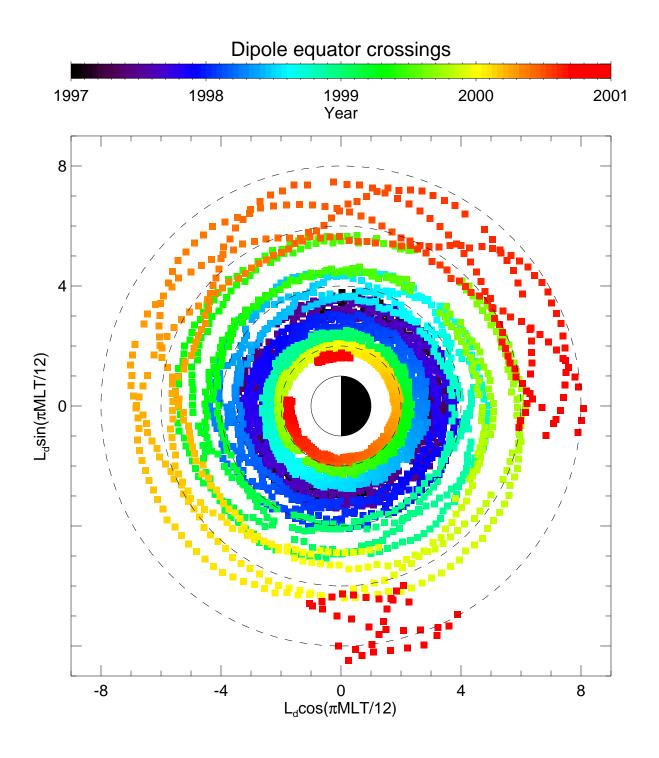
Solar Wind Geospace Pressure Magnetic Fields **Time Time** Global wave periods determined by size of magnetosphere and its internal properties Waves excited by random impulses in solar wind have damped response

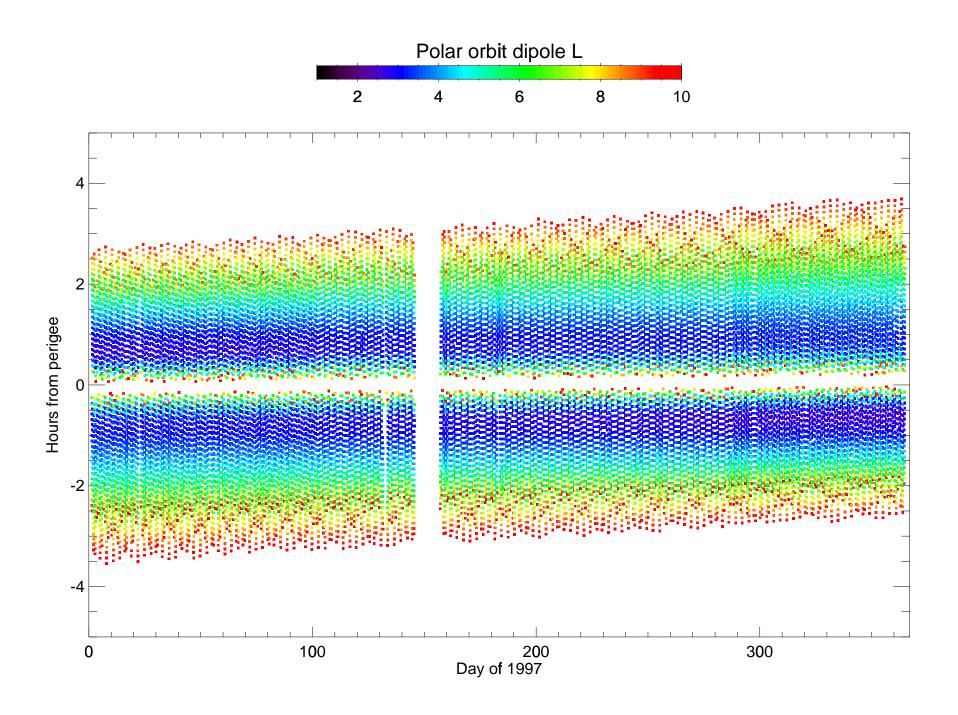
GGS View

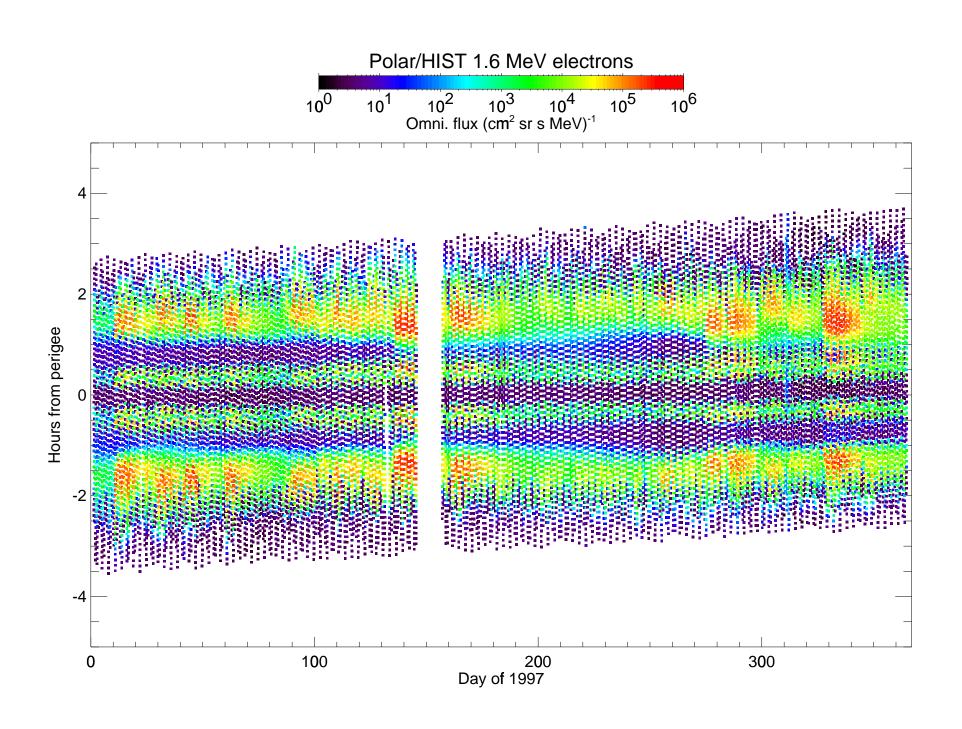


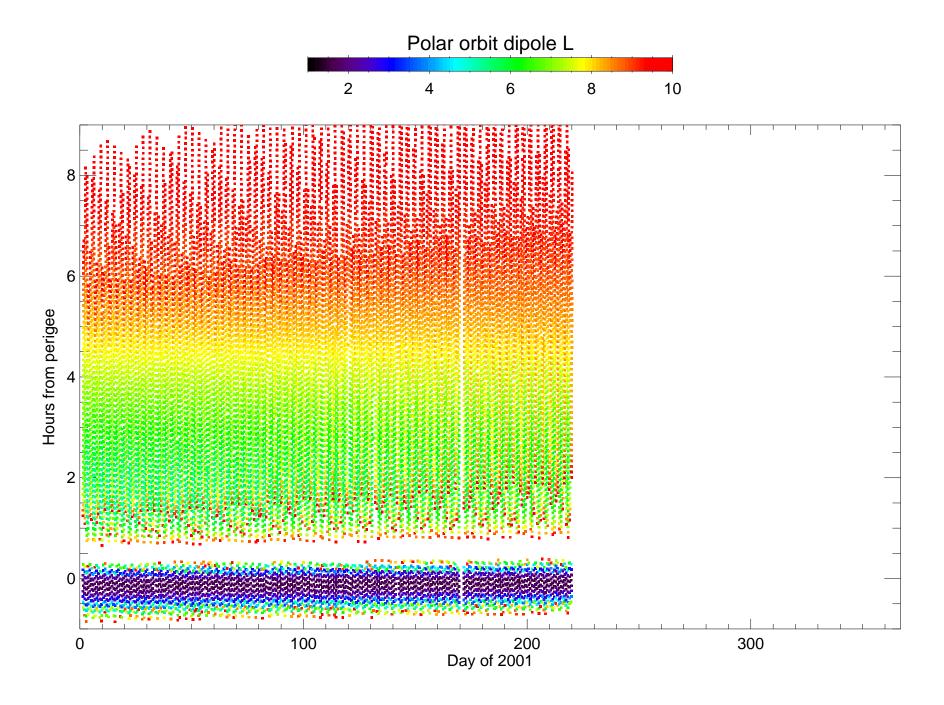


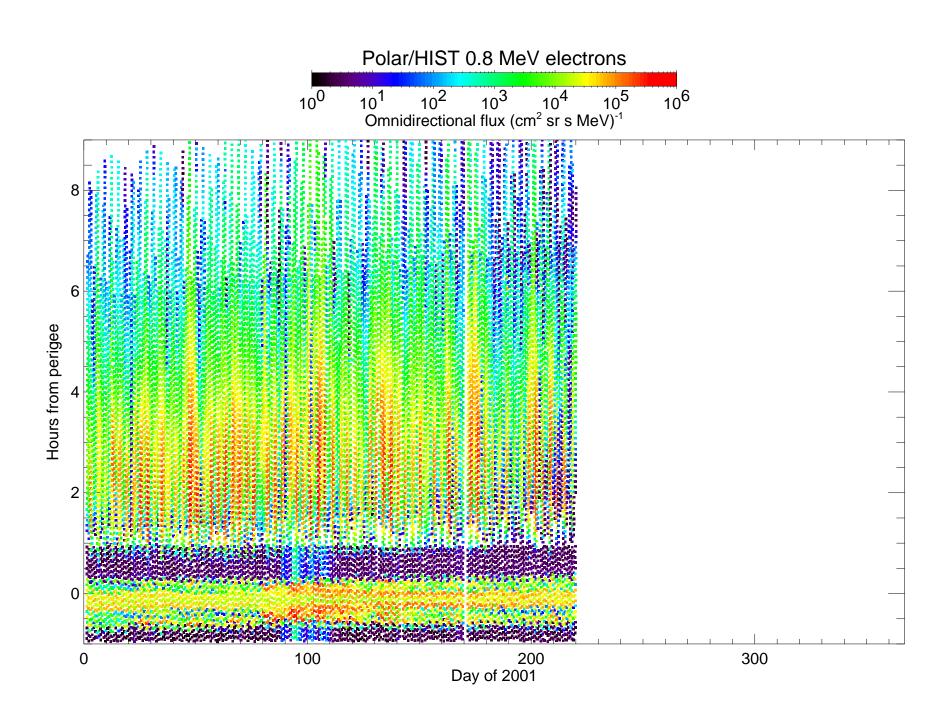
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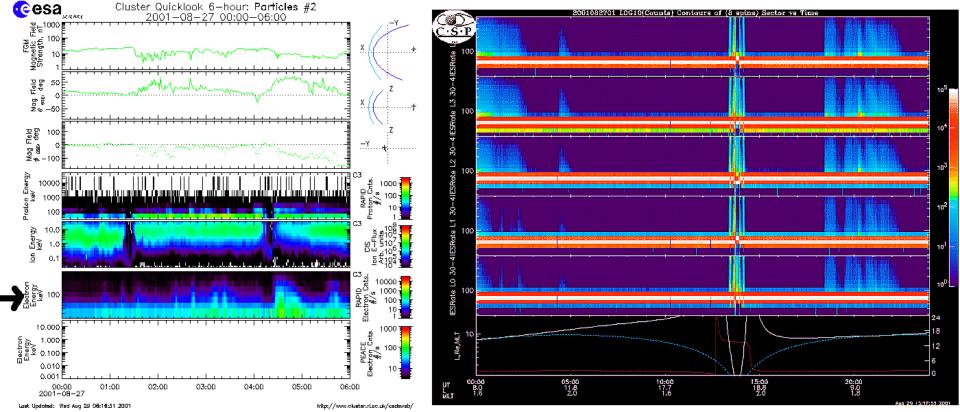












Polar/HIST electrons, 1997-2000, R/L>0.90, 0.0<Kp<6.0

